Scilab Textbook Companion for Introduction To Special Relativity And Space Science by S. P. Singh¹

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July 31, 2019

¹Funded by a grant from the National Mission on Education through ICT, http://spoken-tutorial.org/NMEICT-Intro. This Textbook Companion and Scilab codes written in it can be downloaded from the "Textbook Companion Project" section at the website http://scilab.in

Book Description

Title: Introduction To Special Relativity And Space Science

Author: S. P. Singh

Publisher: Wiley, India

Edition: 1

Year: 2012

ISBN: 9788126532544

Scilab numbering policy used in this document and the relation to the above book.

Exa Example (Solved example)

Eqn Equation (Particular equation of the above book)

AP Appendix to Example(Scilab Code that is an Appednix to a particular Example of the above book)

For example, Exa 3.51 means solved example 3.51 of this book. Sec 2.3 means a scilab code whose theory is explained in Section 2.3 of the book.

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Chapter 1

Interference Diffraction and Polarization

Scilab code Exa 1.1 Determination of slit separation in Young double slit experime

```
1 clc //Given that
2 Beta=0.10//fringe width in cm
3 D=200// separation between source and screen in cm
4 lambda=0.00055// wavelength of incident light in cm
5 //Sample Problem 1 Page No. 46
6 printf ("\n # Problem 1 # \n")
7 d= (D*lambda)/ (10*Beta)
8 printf (" \n Standard formula used \n beta= lambda* D/d \n")
9
10 printf ("\n Separation between sources is %f cm. \n",d)
```

Scilab code Exa 1.2 Determination of wavelength of light in Youngs double slit exp

```
1 clc
2 //Given that
3 D=80// separation between source and screen in cm
4 d=0.18// separation between sources in cm
5 n=4// order of fringe
6 x_n=1.08// distance from central bright fringe in cm
7 //Sample Problem 2 Page No. 47
8 printf("\n # Problem 2 # \n")
9 printf("\n Standard formula used \n x_n= n*lambda *D/d \n")
10
11 lambda=d*x_n/(D*n)*1e7
12 printf("\n Wavelength of light used is %d Angstrom."
, lambda)
```

Scilab code Exa 1.3 Determination of wavelength of light in biprism experiment

```
clc
//Given that
beta=0.0320//fringe width in cm

D=100// separation between source and screen in cm

d=0.184// separation between sources in cm

//Sample Problem 3 Page No. 47

printf ("\n # Problem 3 # \n")

printf("\n Standard formula used beta=lambda*D/d \n")

lambda=d*beta/D*1e8

printf("\n Wavelength of light used is %d Angstrom."
,lambda)
```

Scilab code Exa 1.4 Determination of wavelength of light in biprism experiment

```
1 clc
2 //Given that
3 beta=0.02//fringe width in cm
4 D=100// separation between source and screen in cm
5 u=30// separation between slit and convex lens in cm
6 I=0.7// separation between two images of slits on
      screen in cm
7 //Sample Problem4 Page No. 47
8 printf("\n \# Problem 4 \# \n")
9 printf(" \n Standard formula used \n beta=lambda*D/d
       \n")
10
11 \quad v = 100 - u
12 \ 0 = I * u / v
13 d = 0
14 lambda=d*beta/D*1e8
15 printf("\n Wavelength of light used is %d Angstrom."
      , lambda)
```

Scilab code Exa 1.5 Determination of wavelength of light in biprism experiment

```
beta=0.02//fringe width in cm
b=120// separation between source and eyepiece in cm
d=0.076// separation between sources in cm
//Sample Problem 5 Page No. 47
printf ("\n # Problem 5 # \n")
printf ("\n Standard formula used \n beta= lambda *D/d \n")
beta=x_n/N // calculation of angle formed
lambda=d*beta/D*1e8 // calculation of Wavelength of light
printf ("\n Wavelength of light used is %d Angstrom.", lambda)
```

Scilab code Exa 1.6 Determination of fringe width in biprism experiment

```
1 clc
2 //Given that
3 mu = 1.5 // refractive index of plane glass prism
4 theta = \%pi / 180 // angle of prism
5 y1 = 10 // separation between slit and biprism in cm
6 y2 = 100 //separation sbetween biprism and screen in
  lambda = 0.00005893// wavelength of incident light
     in cm
8 //Sample Problem 6 Page No. 48
9 printf("\n \# Problem 6 \# \n")
10 printf("\n Standard formula used \n Beta = (D *
     lambda) / d")
11 d = 2 * (mu -1) * theta * y1
12 D = y1 + y2
13 Beta = (D * lambda) / d
14 printf("\n Fringe width observed at distance 1 meter
       is %f m", Beta)
```

Scilab code Exa 1.7 Calculation of angle at vertex of biprism

```
1 clc
2 //Given that
3 mu=1.52// refractive index of plane glass prism
4 theta=%pi/180// angle of prism
5 y1=25// separation between slit and biprism in cm
6 y2=175//separation between biprism and screen in cm
7 lambda=0.000055// wavelength of incident light in cm
8 beta=0.02//fringe width in cm
9 //Sample Problem 7 Page No. 49
10 printf("\n # Problem 7 # \n")
11 printf(" \n Standard formula used \n beta= lambda*
     D/d. \ n")
12 D = y1 + y2
13 d = (D*lambda)/beta
14 theta=d/(2*(mu-1)*y1)
15 vertex_angle=180-(2*theta*180/%pi)
16 printf("\n Vertex angle of biprism is %f degree.",
     vertex_angle)
```

Scilab code Exa 1.8 Determination of thickness of glass sheet introduced in Youngs

```
1 clc
2 //Given that
3 mu=1.60// refractive index of plane glass prism
4 lambda=0.0000589// wavelength of incident light in cm
```

```
5 N=15// order of fringe
6 //Sample Problem 8 Page No. 49
7 printf("\n # Problem 8 # \n")
8 printf("\n Standard formula used \n del_x = D/2d
        *(mu-1)*t \n")
9 t=N*lambda/(mu-1)
10 printf("\n Thickness of sheet is %e cm.", t)
```

Scilab code Exa 1.9 Determination of refractive index of glass sheet introduced in

Scilab code Exa 1.10 Determination of lowest order which will be absent from light

```
1 clc
2 //Given that
3 t = 5e-5 // thickness of soap film in cm
4 theta = 35 // angle of view in degree
5 mu = 1.33 // refractive index of soap film
```

```
6 // sample problem 10 page No. 50
7 printf("\n \# Problem 10 \# \n")
9 printf("Standard formula used \n\t 2*mu*t*cos(r) = n
      *lambda ")
10 r = asin (sin(theta * \%pi /180) / mu)
11 \text{ for } n = 1:3
12
       lambda = 2 * mu * t * cos(r) / n
13
       if lambda > t then
            a = a + 1
14
15
       end
16 \, \text{end}
17 printf (" \n The lowest order n = \%d will be
      absent in visible region.",a)
```

Scilab code Exa 1.11 Determination of wavelength of light in biprism experiment

Scilab code Exa 1.12 Determination of wavelength of light in Newtons ring experime

```
1 clc
2 //Given that
3 D15=0.59// diameter of 15th newtons ring in cm
4 D5=0.336// diameter of 5th newtons ring in cm
5 R=100// radius of Plano convex lens in cm
6 //Sample Problem 12 Page No. 51
7 printf("\n # Problem 12 # \n")
8 p=15-5
9 printf("\n Standard formula used \n D_a^2 D_b^2
= 4*p*R*lambda \n")
10
11 lambda=(D15^2 - D5^2)/(4*p*R)*1e8
12 printf("\n Wavelength of light used is %d Angstrom."
, lambda)
```

Scilab code Exa 1.13 Calculation of refractive index of air

```
1 clc
2 //Given
3 t=40// length of tube in cm
4 lambda=5e-5// wavelength of incident light in cm
5 n=150// order of fringe
6 //Sample Problem 13 Page No. 52
7 printf("\n # Problem 13 # \n")
8 printf(" \n Standard formula used \n (mu 1)*t
= n* lambda \n")
```

```
9 t=n*lambda/t+1 10 printf("\n Refractive index of oil film is %f .", t )
```

Scilab code Exa 1.14 Calculation of thickness of air film

Scilab code Exa 1.15 Calculation of thickness of oil film

```
1 clc
2 //Given
3 mu_oil=1.3// refractive index of oil
4 mu_glass=1.5//refractive index of glass
5 lambda1=5e-7// wavelength of incident light in cm
```

 ${
m Scilab\ code\ Exa\ 1.16}$ Determination of diameter fourth bright ring in Newtons ring

Scilab code Exa 1.17 Determination of radius of curvature of Plano convex lens use

Scilab code Exa 1.18.1 Determination of radius of curvature of Plano convex lens u

```
1 clc
2 //Given
3 D_10=0.5// diameter of 10th dark ring
4 lambda=5.9e-5// wavelength of incident light in cm
5 n=10// order of ring
6 //Sample Problem 18a Page No. 54
7 printf("\n # Problem 18a # \n")
8 printf("\n Standard formula used \n r_n^2 = n* lambda*R \n")
9 r=D_10/2
10 R=r^2/ (n*lambda)/1000
11 printf("Radius of curvature is %f m.\n ",R)
```

Scilab code Exa 1.18.2 Determination of thickness of air film in Newtons ring expe

```
1 clc
2 //Given
3 lambda=5.9e-5// wavelength of incident light in cm
4 n=10// order of ring
5 //Sample Problem 18b Page No. 54
6 printf("\n # Problem 18b # \n")
7 printf(" \n Standard formula used \n 2t = n*lambda \n")
8 t=n*lambda/200
9 printf(" Thickness of air film is %e m.\n ",t)
```

Scilab code Exa 1.19 Determination of radius of curvature of Plano convex lens use

```
1 clc
2 //Given
3 mu=4/3
4 D_10=0.6// diameter of tenth ring in cm
5 lambda=6.0e-5// wavelength of incident light in cm
6 n=10// order of ring
7
8 //Sample Problem 19 Page No. 54
9 printf("\n # Problem 19 # \n")
10 printf(" \n Standard formula used \n D_n^2 = 4*n* lambda*R/mu \n")
11 R= (mu*D_10^2/ (4*n*lambda))
12
13 printf(" Radius of curvature of lens is %d cm.\n", ceil(R))
```

Scilab code Exa 1.20 Determination of wavelength of light using grating

```
clc
//Given
grating_element=6000// lines per centimeter
theta=30// angle of second order spectral line in
    degree
n=2// order

//Sample Problem 20 Page No. 54
printf("\n # Problem 20 # \n")
printf(" \n Standard formula used \n n*lambda= (a+b)*sin(theta) \n")

lambda=sin(theta*%pi/180)/(grating_element*n)
printf(" Wavelength is %e cm.\n",lambda)
```

 ${
m Scilab\ code\ Exa\ 1.21}$ Determination of highest order which can be seen by a grating

```
1 clc
2 //GivenS
3 lambda=6.2e-5// wavelength of monochromatic light in cm
4 grating_element= 1/5000// lines per centimeter
5 theta=90// angle of second order spectral line in degree
6
7 //Sample Problem 21 Page No. 55
```

```
8 printf("\n # Problem 21 # \n")
9 printf(" \n Standard formula used \n n*lambda= (a+b)*sin(theta) \n")
10 n=grating_element/lambda
11 printf(" Maximum order n = %d may be seen in between the given wavelength spectrum.\n",n)
```

Scilab code Exa 1.22 Determination of dispersive power grating in third order spec

```
1 clc
2 //Given
3 lambda=5.5e-5// wavelength of monochromatic light in cm
4 grating_element=1/4000// lines per centimeter
5 n=3// order of spectrum
6
7 //Sample Problem 22 Page No. 55
8 printf("\n # Problem 22 # \n")
9 printf(" \n Standard formula used \n n*lambda= (a+b)*sin(theta)\n")
10 sin_theta=n*lambda/grating_element
11 cos_theta=sqrt(1-sin_theta^2)
12 disp_pow=n/ (grating_element*cos_theta)
13 printf (" Dispersive power is %d. \n ",disp_pow)
```

Scilab code Exa 1.23.1 Determination of clearly observation between two wavelength

```
1 clc
2 // Given That
```

```
3 \quad lambda1=5.89e-5// \quad wavelength \quad in \quad cm
4 lambda2=5.896e-5//wavelength in cm
5 n=1// for second order spectrum
6 t = 2 // width of detraction grating
7 grating_element = 425 // no. of lines per cm
8 //Sample Problem 23a Page No. 56
10 printf("\n \# Problem 23a \# \n")
11 printf(" \n Standard formula used \n
                                             lambda /
      d_{\text{lambda}} = n*N \setminus n")
12 total_line = t * grating_element
13 printf("\n Total number of lines on diffraction
      grating is %d \ n \ So, total_line)
14 N=lambda1/ (lambda2-lambda1)/n
15 if (N > total_line) then
16 printf ("\n Lines will not be resolved in %d order
      .",n)
17 printf("\nas %d lines are required for diffraction
       ", N)
18 else printf("\nas %d lines are required
      diffraction are. ", N)
       printf (" Lines will be resolved in %d order",
19
          n)
20 end
```

Scilab code Exa 1.23.2 Determination of clearly observation between two wavelength

```
1 clc
2 // Given That
3 lambda1=5.89e-5// wavelength in cm
4 lambda2=5.896e-5//wavelength in cm
5 n=2// for second order spectrum
6 t = 2 // width of diffraction grating
```

```
7 grating_element = 425 // no. of lines per cm
8 //Sample Problem 23b Page No. 56
10 printf("\n \# Problem 23b \# \n")
11 printf(" \n Standard formula used \n lambda /
     d_{lambda} = n*N \ n")
12 total_line = t * grating_element
13 printf("\n Total number of lines on diffraction
     grating %d \n \n So", total_line)
14 N=lambda1/ (lambda2-lambda1)/n
15 if (N > total_line) then
16 printf ("\n , Lines will not be resolved in %d order
     .",n)
17 printf("\nas %d lines are required for diffraction
      are ", N)
18 else printf (" Lines will be resolved in %d order",
      n)
       printf("\nas %d lines are required
19
          diffraction are . ", N)
20 end
```

 ${f Scilab\ code\ Exa\ 1.24}$ Calculation of maximum number of lines two just resolve two w

```
1 clc
2 // Given That
3 lambda1=5.89e-5// wavelength in cm
4 lambda2=5.896e-5//wavelength in cm
5 t=2.5// width of grating in cm
6 n=2// for second order spectrum
7 //Sample Problem 24 Page No. 56
8
9 printf("\n # Problem 24 # \n")
10 printf(" \n Standard formula used \n lambda /
```

 ${
m Scilab\ code\ Exa\ 1.25}$ Calculation of maximum number of lines two just resolve two w

```
1 clc
2 // Given That
3 a=12e-5// slit width in cm
4 lambda1=5.89e-5// wavelength in cm
5 lambda2=5.896e-5//wavelength in cm
6 n=2// for second order spectrum
7 //Sample Problem 25 Page No. 56
8 printf("\n # Problem 25 # \n")
9 printf(" \n Standard formula used \n lambda / d_lambda = n*N \n\n")
10 d_lambda = lambda2-lambda1
11 grating_element= lambda1/ (d_lambda*n)
12
13 printf(" Minimum number of lines required is %d . \n ", ceil(grating_element))
```

Scilab code Exa 1.26 Determination of half angular width of central maxima in Frau

```
1 clc
2 // Given That
3 a = 12e-5 // slit width in cm
```

Scilab code Exa 1.27 Calculation of resolving power of grating in first order

```
1 clc
2 // Given That
3 \quad lambda1 = 5.9e-5 // wavelength in cm
4 lambda2 = 5.896e-5 //wavelength in cm
5 \text{ lambda} = 5.89e-5 // \text{ wavelength in cm}
6 grating_element = 4000 // lines per cm
7 t = 4 // width of grating in cm
8 n = 1 // \text{ for first order spectrum}
9 //Sample Problem 27 Page No. 58
10 printf("\n \# Problem 27 \# \n")
11 printf(" \n Standard formula used \n lambda /
      d_{lambda} = n * N ) \ n")
12
13 N = t * grating_element
    Resolv_pow = lambda /(lambda2 - lambda)
15
    N = Resolv_pow / n
16
    if (grating_element > N ) then
17
        printf("Grating will well resolve two spectral
18
           lines. \n")
19
        end
```

Scilab code Exa 1.28 Calculation of distance between centre and first fringe

 ${
m Scilab~code~Exa~1.29}$ Comparison of intensities of ordinary and extra ordinary light

```
1 clc
2 // Given That
3 theta = 60 // angle between plane of vibration of incident beam with optic axis
4 
5 //Sample Problem 29 Page No. 859
6 printf("\n # Problem 29 # \n ")
7 printf("Standard formula used is \n I = A^2*cos^2(theta) \n")
```

```
8 ratio = (tan(theta*%pi /180))^2  // ratio of
        extraordinary and aordinary intensites
9 printf("Ratio of extraordinary and ordinary
        intensites is %f .", ratio)
```

Scilab code Exa 1.30 Calculation of thickness of half wave plate

```
1 clc
2 // Given That
3 mu_e = 1.553 // refractive index of quartz plate for
        extra ordinary light
4 mu_o = 1.544 // refractive index of quartz plate for
        ordinary light
5 lambda = 5.89e-5 // wavelength of light in Angstrom.
6 //Sample Problem 30 Page No. 859
7 printf("\n # Problem 30 # \n ")
8 printf("Standard formula used is \n lambda= 2t(mu_e-mu_o) \n")
9 t = lambda / (2 * (mu_e - mu_o))
10 printf("Thickness of half wave plate of quartz is %e cm.", t)
```

Scilab code Exa 1.31 Determination of least thickness to get a plane polarized lig

Scilab code Exa 1.32 Calculation of difference between refractive indices

```
clc
// Given That
lambda=5.89e-5// wavelength in cm
rotation=(%pi/18)// rotation of plane of
    polarization in degree per cm

//Sample Problem 32 Page No. 60
printf("\n # Problem 32 # \n ")
printf(" \n Standard formula used \n delta=pi*d*
    del_mu/lambda \n")
del_mu=rotation*lambda/ (%pi)
printf("Difference in refractive indices of substance is %e .\n",del_mu)
```

Scilab code Exa 1.33 Determination of specific rotation of sugar solution

1 clc

Chapter 2

Electromagnetic Waves

Scilab code Exa 2.2 Calculation of energy stored in cylinder and wave Intensity

```
1 clc
2 // Given That
3 E_0 = 50 // magnitude of electric field in N/C
4 l = 100 // radius of cylinder in cm
5 a = 5 // area of cross section in cm<sup>2</sup>
6 c = 3e8 // speed of light in m/s
7 epsilon_0 = 8.85e-12 // permittivity of free space
8 //Sample Problem 2 Page No. 79
9 printf("\n \# Problem 2 \# \n ")
10 v = a*1e-4 *1*1e-2//calculation of volume of
     cylinder
11 u = (1/2)*epsilon_0*E_0^2/calculation of energy
     intensity
12 U = v*u//calculation of Energy contained in cylinder
13 I = u*c//calculation of Intensity of wave
14 printf ("Energy contained in cylinder is \%eJ \ n", U)
15 printf ("Intensity of wave is \%fW/m^2", I)
```

Scilab code Exa 2.3 Determination of amplitude of electric field

```
1 clc
2 // Given That
3 I = 2.4 // intensity of radiation in Watt per meter square
4 epsilon_0 = 8.85e-12
5 c = 3e8
6 //Sample Problem 3 Page No. 80
7 printf("\n # Problem 3 # \n ")
8 E = sqrt ((2* I)/ (c * epsilon_0)) // calculation of amplitude of electric field is
9 printf("Amplitude of electric field is %f N/C \n", E
)
```

Scilab code Exa 2.4 Calculation of energy stored in a length of laser beam

```
1 clc
2 // Given That
3 l = 75 // length of laser beam in cm
4 power = 6e-3 // power of beam in mW
5 c = 3e8
6 //Sample Problem 4 Page No. 80
7 printf("\n # Problem 4 # \n ")
8 t = l / (c * 100) //calculation of time taken to cover distance
9 U = power/1000 * t//calculation of Energy stored in given length
```

```
10 \operatorname{printf}("Energy stored in given length is \%e J \n", U)
```

Scilab code Exa 2.6 Calculation of maximum Electric and magnetic force on an elect

```
1 clc
2 // Given That
3 E_0 = 300 // \text{maximum electric field in}
     electromagnetic wave in w/m
4 v = 2e8 // \text{ speed of moving electron in m/s along y} -
      axis
5 c = 3e8 // speed of light in m/s
6 q = 1.6e-19 // charge on electron in coulomb
7 //Sample Problem 6 Page No. 81
8 printf ("\n \# Problem 6 \# \n ")
9 B_0 = E_0 / c // calculation of magnitude of
     maximum magnetic field
10 F_e = q*E_0 // calculation of electromagnetic force
     on electron in N
11 F_b = q*v*B_0 // calculation of magnetic force on
     electron in N
12
   printf ("The maximum electric force on electron is
       \%e N along y -axis \n", F_e)
   printf("The maximum magnetic force on electron is
13
      \%e N along z - axis\n", F_b)
```

Scilab code Exa 2.7 Calculation of average solar energy flux and pressure applied

```
1 clc
```

```
2 // Given That
3 d = 1.5e11 // separation between earth and sun in
    meter
4 power_sun = 3.8e26// power radiated by sun in W
5 c = 3e8
6 //Sample Problem 7 Page No. 82
7 printf("\n # Problem 7 # \n ")
8 s = power_sun /(4 * %pi * (d^2)) //calculation of
    Energy received per unit surface area per unit
    time
9 p = s / c // calculation of Pressure applied by sun
    radiations on earth
10 printf("Energy received per unit surface area per
    unit time is %f", s)
11 printf("\n Pressure applied by sun radiations on
    earth is %e N/m^2 \n", p)
```

Scilab code Exa 2.8 Calculation of electric flux through plane square due to unifo

Scilab code Exa 2.9 Calculation of electric flux through plane circle due to unifo

 ${
m Scilab\ code\ Exa\ 2.10}$ Determination of charge inside sphere with the help of electrons

```
1 clc
2 // Given That
3 A = 200 // magnitude of electric field in V/m^2
4 epsilon_0 = 8.85e-12 // permittivity of free space
5 a = 20 // radius of sphere in cm
6 //Sample Problem 10 Page No. 84
7 printf("\n # Problem 10 # \n")
```

Scilab code Exa 2.11 Calculation of induced emf current due to changing area and

```
1 clc
2 // Given That
3 B = 0.2 // magnetic field in T
4 del_r = 1 // rate of change of decrement in loop
    radius in cm/s
5 r = 20 // radius of frame in cm
6 R = 10 // resistance of frame in m ohm
7 //Sample Problem 11 Page No. 84
8 printf("\n # Problem 11 # \n ")
9 e = 2* %pi * B *r *1e-2* del_r*1e-2 // magnitude of
    emf induced in coil
10 i = (e) / (R*1e-3) //calculation of Current induced
    due to changing magnetic field
11 printf("Current induced due to changing magnetic
    field is %f A \n", i)
```

Scilab code Exa 2.12 Calculation of induced current due to changing magnetic field

```
1 clc
2 // Given That
3 phi = 0.02 // rate of change of magnetic field in T/
s
```

```
4 r = 2 // radius of frame in cm
5 R = 2 // resistance of frame in m ohm
6 //Sample Problem 12 Page No. 85
7 printf("\n # Problem 12 # \n ")
8 a = %pi * (r*1e-2)^2
9 e = a * phi // magnitude of emf induced in coil
10 i = (e) / (R*1e-3)
11 printf("Current induced due to changing magnetic field is %f mA \n", i*1000)
```

Scilab code Exa 2.13 Calculation of poynting vector at the surface of sun

```
1 clc
2 // Given That
3 r = 7e8 // radius sun in meter
4 power_sun = 3.8e26// power radiated by sun in W
5 //Sample Problem 13 Page No. 86
6 printf("\n # Problem 13 # \n ")
7 s = power_sun /(4 * %pi * (r^2)) //calculation of
    Pressure applied by sun radiations on earth
8 printf("Pressure applied by sun radiations on earth
    is %e W/m^2 \n", s)
```

 ${
m Scilab\ code\ Exa\ 2.14\ Calculation\ of\ amplitude\ of\ electric\ and\ magnetic\ fields\ of\ s}$

```
1 clc
2 // Given That
3 solar_const = 2 // energy received by earth from sun
      in Cal/min cm2
```

```
4 mu_not = 1.2566e-6 // universal constant
5 epsilon_not = 8.85e-12 // universal constant
6 //Sample Problem 14 Page No. 86
7 printf("\n # Problem 14 # \n ")
8 ratio = sqrt(mu_not / epsilon_not) // constant
9 E = sqrt (ratio *4.2 * solar_const / 6e-3)
10 E_not = E * sqrt(2) //calculation of Amplitude of electric vectors
11 H_not = E_not / ratio//calculation of Amplitude of magnetic vectors
12 printf("Amplitude of electrical and magnetic vectors are given as %f V/m and %f A/m", E_not, H_not)
```

 ${
m Scilab~code~Exa~2.15}$ Calculation of average value of intensity of electric field r

```
1 clc
2 // Given That
3 r = 1 // distance from lamp in meter
4 power = 100// power radiated by lamp in W
5 mu_not = 1.2566e-6 // universal constant
6 epsilon_not = 8.85e-12 // universal constant
7 //Sample Problem 15 Page No. 87
8 printf("\n # Problem 15 # \n ")
9 s = power /(4 * \%pi * (r^2)) // calculation of
     intensity at a distance
10 ratio = sqrt(mu_not / epsilon_not) //calculation of
     a constant
11 E = sqrt (ratio * s) //calculation of Average value
     of intensity of electric field
12 printf("Average value of intensity of electric
      field is \%f V/m \n", E)
```

Chapter 3

Dual Nature of Light

Scilab code Exa 3.1 Calculation of velocity of ejected photoelectrons

```
1 clc
2 //Given that
3 h = 6.6e-34 // plank's constant
4 nu = 2e15 // frequency in Hz
5 phi = 6.72e-19
6 m = 9e-31
7 //Sample Problem 1 Page No. 135
8 printf("\n\n\n # Problem 1 # \n")
9 printf("Standard formula Used \n ( 1/2)*m*v^2 = h*nu - phi")
10 v = sqrt ((h * nu)/ m) //calculation of maximum velocity of photoelectron
11 printf("\n Maximum velocity of photoelectron can be %e m/s..", v)
```

Scilab code Exa 3.2 Calculation of energy of ejected photoelectrons

```
1 clc
2 //Given that
3 h = 6.6e-34 // plank's constant
4 lambda_threshold = 2.4e-7 // threshold wavelength in
5 lambda = 2e-7 // wavelength of irradicated light in
     photo emmission
6 c = 3e8
7 //Sample Problem 2 Page No. 135
8 printf("\n \# Problem 2 \# \n")
9 printf("\n Standard formula Used \n E = h * (nu1)
      nu2)")
10 E = h * c * ((lambda_threshold - lambda)/(lambda *
     lambda_threshold))/1.6e-19 // calculation of
     nergy of photoelectrons
11 printf("\n Energy of photoelectrons emitted is %f
     \mathrm{eV}", E)
```

 ${f Scilab\ code\ Exa\ 3.3}$ Determination of shortest wavelength emitted in X ray operation

```
1 clc
2 //Given that
3 applied_voltage = 4e4 // in volt
4 h = 6.624e-34 // plank's constant
5 c = 3e8 // speed of light
6 e = 1.6e-19 // charge on electron
7 //Sample Problem 3 Page No. 136
8 printf("\n\n\n # Problem 3 # \n")
9 printf("\n Standard formula Used \n E = h*c/lambda")
10 lambda = h * c / ( e * applied_voltage) *1e10 // calculation of Shortest wavelength emitted
```

```
11 printf("\n Shortest wavelength emitted is %f Angstrom.", lambda)
```

Scilab code Exa 3.4 Calculation of velocity of a moving electrons

Scilab code Exa 3.5 Determination of threshold wavelength for photoemission

```
1 clc
2 //Given that
3 phi = 6 // work function in eV
4 h = 6.624e-34 // plank's constant
5 c = 3e8 // speed of light
6 e = 1.6e-19 // charge on electron
```

Scilab code Exa 3.6 Determination of Compton shift of X ray photon

```
1 clc
2 //Given that
3 theta = %pi/2 // scattering angle of photon
4 h = 6.624e-34 // plank's constant
5 c = 3e8 // speed of light
6 e = 1.6e-19 // charge on electron in coloumb
7 m_e = 9.1e-31 // mass of electron in kg
8 //Sample Problem 6 Page No. 137
9 printf("\n\n\n # Problem 6 # \n")
10 printf("\n Standard formula Used \n delta_lambda = h
          * (1 - cos (theta )) / ( m_e * c)")
11 delta_lambda = h * (1 - cos (theta )) /( m_e * c) //
          calculation of Change in wavelength of electron
12 printf("\n Change in wavelength of electron is %f
          Angstrom.", delta_lambda*1e10)
```

Scilab code Exa 3.7 Calculation of de Broglie wavelength of moving particle

```
1 clc
2 //Given that
3 angle = \%pi/2 // scattering angle of photon
4 h = 6.624e-34 // plank's constant
5 v = 2e6 // speed of particle
6 = 1.6e-19 // charge on electron
7 m = 1e-3 // mass of particle in kg
8 //Sample Problem 7 Page No. 137
9 printf("\n\n" Problem 7 # \n")
10 printf("\n Standard formula Used \n lambda = h / (m
     * v)")
11 lambda = h / (m * v) //calculation of de Broglie
     wavelength of particle
12 printf("\n de Broglie wavelength of particle is %e
     m.", lambda)
13 printf("\n Here the de Broglie wavelength is too
     small to be detected. This wavelength is far
     smaller than the wavelength of X ray.\n Hence
      diffraction experiment with such a stream of
      particle will not be successful.")
```

Scilab code Exa 3.8.1 Determination of possibility of photo emission and velocity

```
1 clc
2 //Given that
3 lambda = 4.3e-7 // wavelength of light in meter
4 phi_Ni = 5 // work function of nickel in eV
5 h = 6.624e-34 // plank's constant
6 c = 3e8 // speed of light
7 m_e = 9.1e-31 // mass of electron in kg
8 //Sample Problem 8a Page No. 138
9 printf("\n\n\n # Problem 8a # \n")
10 lambda_threshold = h * c / (phi_Ni*1e-19) //
```

```
calculation of longest wavelength required
11 if (lambda_threshold < lambda) then
       printf("\n As the threshold wavelength is less
12
          than wavelength of incident radiation \n So
          electron will not be ejected \n")
13 else
14
      v = sqrt((2* h * c *(lambda - lambda_threshold))
          / (m * lambda_threshold * lambda ))) //
          calculation of ejected velocity Electron
       printf("\n As the threshold wavelength is
15
          greater than wavelength of incident radiation
          So electron will be ejected with velocity %e
          . ",v)
16 end
```

Scilab code Exa 3.8.2 Determination of possibility of photo emission and velocity

```
1 clc
2 //Given that
3 lambda = 4.3e-7 // wavelength of light in meter
4 phi_K = 2.3 // work function of nickel in eV
5 h = 6.624e-34 // plank's constant
6 c = 3e8 // speed of light
7 m_e = 9.1e-31 // mass of electron in kg
8 //Sample Problem 8b Page No. 138
9 printf("\n\n" | Problem 8b # \n")
10 lambda_threshold = h * c / (phi_K *1.6e-19) //
     calculation of longest wavelength required
11 if (lambda_threshold < lambda) then
       printf ("As the threshold wavelength is less than
12
           wavelength of incident radiation Solectron
          will not be ejected \n")
13 else
```

 ${
m Scilab\ code\ Exa\ 3.9}$ Determination of wavelength for which the second order Bragg r

Scilab code Exa 3.10.1 Calculation of Inter atomic separation

Scilab code Exa 3.10.2 Calculation of angle for secondary maxima

```
1 clc
2 //Given that
3 n = 2 // order
4 lambda = 5.2e-11 // wavelength in Angstrom
5 d = 2.98e-10 // interatomic separation in Angstrom
6 //Sample Problem 10b page No. 139
7 printf("\n\n" Problem 10b # \n")
8 printf("\n Standard formula Used \n 2 * d * sin(
     theta) = n * lambda ")
9 theta_rad = asin ( (n * lambda) / ( 2 * d)) //
     calculation of angle for secondary maxima in
     radian
10 theta_deg = theta_rad * 180 / \%pi //calculation of
     angle for secondary maxima in degree
11 printf ("\n Angle for secondary maxima is
     theta_deg )
```

Scilab code Exa 3.11 Calculation of X ray frequency after scattering through 90 de

```
1 clc
2 //Given that
3 \text{ nu} = 3.2e19 // \text{frequency in hartz}
4 theta = 90 // angle of scattered photon in degree
5 \text{ m_e} = 9.1\text{e}-31 \text{ // mass of electron in Kg}
6 c = 3e8 // speed of light in m/s
7 h = 6.626e-34 // plank's constant
8 //Sample Problem 11 Page No. 140
9 printf("\n \n\n Problem 11 # \n")
10 printf("\n Standard formula Used \n delta_lambda = h
      * (1 - \cos (theta)) / (m_e * c)")
11 lambda = c / nu / calculation of incident wavelength
12 lambda_shift = h *(1 - cos(theta * \%pi / 180))/(
     m_e * c) //calculation of shift in wavelength
13 lambda1 = lambda + lambda_shift//calculation of
      wavelength of scattered photon
14 nu1 = c / lambda1 // calculation of Frequency after
      scattering
15 printf ("\n Frequency after scattering is %e Hz.",
      nu1)
```

Scilab code Exa 3.12 Calculation of uncertainty inn momentum of electron if it is

```
1 clc
2 //Given that
```

```
3 r = 1e-14 // radius of nucleus of atom in meter
4 h = 6.626e-34 // Plank's constant
5 //Sample Problem 12 page No. 140
6 printf("\n\n\n # Problem 12 # \n")
7 printf("\n Standard formula Used \n delta_p *
        delta_x >= h /(2*pi)")
8 del_x = 2 * r //calculation of Uncertainty in
        position
9 del_p = h / (2 * %pi * del_x) //calculation of
        Uncertainty in momentum
10 printf ("\n Uncertainty in momentum is %e Kg-m/s.",
        del_p)
```

Scilab code Exa 3.13 Calculation of uncertainty in position of a moving electron

```
1 clc
2 //Given that
3 v = 300 // speed of electron in m/s
4 accuracy = 1e-4 // accuracy in speed
5 h = 6.6e-34 // Plank's constant
6 \text{ m_e} = 9.1\text{e}-31 \text{ // mass of electron in Kg}
7 //Sample Problem 13 page No. 140
8 printf("\n\n" + Problem 13 # \n")
9 printf("\n Standard formula Used \n delta_p *
      delta_x >= h /(2*pi)")
10 del_p = accuracy * m_e * v //calculation of
      Uncertainty in momentum
11 del_x = h / (4 * \%pi * del_p) // calculation of
      Uncertainty in position
12 printf ("\n Uncertainty in position of electron is
     \% f mm. ", del_x*1000 )
```

Scilab code Exa 3.14 Calculation of wavelength of first spectral line of Lyman ser

```
1 clc
2 //Given that
3 lambda1 = 6560 // wavelength in Angstrom
4 n1 = 1 // transition state no
5 n2 = 2 // transition state no
6 \text{ n3} = 3 \text{ // transition state no.}
7 //Sample Problem 14 page No. 141
8 printf("\n\n" + Problem 14 # \n")
9 printf("\n Standard formula Used \n\n For Balmer
     Series \n 1/\lambda = R*(1-(1/n)^2) \n\n For
     Lyman series \n 1/\lambda = R*((1/2)^2 - (1/n)^2)"
10 lambda2 = (n2^2 * n1^2) * (n3^2 - n2^2) / ((n2^2 - n1))
     ^{2}) * (n3^{2} * n2^{2})) * lambda1 //calculation of
     Wavelength of first line of Lyman series
11 printf ("\n \nWavelength of first line of Lyman
      series is %f Angstrom. ", lambda2 )
```

Scilab code Exa 3.15 Calculation of zero energy of a linear harmonic oscillator

```
1 clc
2 //Given that
3 m = 2e-3 // mass of linear harmonic oscillator in kg
4 k = 100 // spring constant in N/m
5 h = 6.6e-34 // Plank's constant
6 //Sample Problem 15 page No. 142
```

```
7 printf("\n\n\ # Problem 15 # \n")
8 printf("\n Standard formula Used \n f = sqrt(k / m
      ) \n U = 1/2* h * nu ")
9 nu = sqrt(k / m ) / (2 * %pi) //calculation of
      frequency of linear harmonic oscillator
10 U = 1/2* h * nu //calculation of Zero point energy
      of a linear harmonic oscillator
11 printf ("\n Zero point energy of a linear harmonic
      oscillator is %e J.", U )
```

 ${f Scilab\ code\ Exa\ 3.16.1}$ Calculation of wavelength of first spectral line of Lyman s

```
1 clc
2 //Given that
3 R = 1.097 // Rydbergs constant
4 n1 = 1 // transition state no
5 n2 = 2 // transition state no
6 //Sample Problem 16a page No. 142
7 printf("\n\n" Problem 16a # \n")
8 printf ("\n Standard formula Used \n For Lyman
     series 1/\text{lambda} = R*((1/2) ^2 - (1/n) ^2)")
9 nu1 = R * (n2^2 - n1^2) / (n1^2 * n2^2) //
     calculation of frequency of first line of Lyman
     series
10 lambda1 = 1/ nu1//calculation of Wavelength of first
      line of Lyman series
11 printf ("\n Wavelength of first line of Lyman series
      is %f Angstrom. ", lambda1 *1000)
```

Scilab code Exa 3.16.2 Calculation of wavelength of second spectral line of Lyman

```
1 clc
2 //Given that
3 R = 1.097 // Rydbergs constant
4 n1 = 1 // transition state no
5 n2 = 3 // transition state no
6 //Sample Problem 16b page No. 142
7 printf("\n\n # Problem 16b # \n")
8 printf("\n Standard formula Used \n For Lyman series
      1/lambda = R*((1/2)^2 -(1/n)^2)
9 \text{ nu1} = R * (n2^2 - n1^2) / (n1^2 * n2^2) //
     calculation of frequency of first line of Lyman
     series
10 lambda1 = 1/ nu1 //calculation of Wavelength of
     first line of Lyman series
11 printf ("\n Wavelength of second line of Lyman
     series is %d Angstrom. ", lambda1 *1000 )
```

Scilab code Exa 3.17 Calculation of temperature at which it will emit a wavelength

```
1 clc
2 //Given that
3 lambda1 = 4700 // wavelength in Angstrom
4 lambda2 = 1.4e-5//wavelength in cm
5 temp1 = 6174 // temperature of a black of in kelvin
6 //Sample Problem 17 page No. 143
7 printf("\n\n\n # Problem 17 # \n")
8 printf("\n Standard formula Used \n lambda * T = constant")
9 temp2 = lambda1 * temp1 / (lambda2 * 1e8) // calculation of temperature
10 printf ("\n Blackbody will emit wavelength 1.4e-5
```

Scilab code Exa 3.19.1 Calculation of Compton shift for 90 degree scattering

```
1 clc
2 //Given that
3 lambda = 1 // wavelength in Angstrom
4 theta = 90 // angle of scattered photon in degree
5 m_e = 9.11e-31 // mass of electron in Kg
6 c = 3e8 // speed of light in m/s
7 h = 6.63e-34 // plank's constant
8 //Sample Problem 19a page No. 144
9 printf("\n\n\n # Problem 19a # \n")
10 printf("\n Standard formula Used \n delta_lambda = h
          * (1 - cos (theta )) / (m_e * c)")
11 lambda_shift = h *(1 - cos(theta * %pi / 180))/(
          m_e * c) //calculation of Change in frequency
12 printf ("\n Change in frequency is %f Hz.",
          lambda_shift * 1e10)
```

 ${f Scilab\ code\ Exa\ 3.19.2}$ Calculation of kinetic energy imparted to electron in Compt

```
1 clc
2 //Given that
3 lambda1 = 1 // wavelength in Angstrom
4 lambda2 = 1.0243 // wavelength in Angstrom
5 c = 3e8 // speed of light in m/s
6 h = 6.63e-34 // plank's constant
7 //Sample Problem 19b page No. 144
```

```
8 printf("\n\n\n # Problem 19b # \n")
9 printf("\n Standard formula Used \n E= h *(nu1 nu2)")
10 K = h * c * (( lambda2 - lambda1 )/ (lambda1 * lambda2 )) *(10e9 / 1.6e-19) //calculation of Kinetic energy imparted to recoiling
11 printf ("\n Kinetic energy imparted to recoiling electron is %d eV.", K)
```

Scilab code Exa 3.20 Calculation of final wavelength of scattered photon in Compto

```
1 clc
2 //Given that
3 theta = 90 // angle of scattered photon in degree
4 E_rest = 938.3 // rest mass energy of a proton in
5 E = 12// \text{ energy of scattered proton in Mev}
6 c = 3e8 // speed of light in m/s
7 h = 6.63e-34 // plank's constant
8 //Sample Problem 20 page No. 145
9 printf("\n\n" + Problem 20 # \n")
10 printf("\n Standard formula Used \n delta_lambda = h
      * (1 - \cos (theta)) / (m_e * c)")
11 lambda = h * c / (E * 1.6e-13) // calculation of
     incident wavelength
12 lambda1 = lambda + h * c / (E_rest * 1.6e-13)
      calculation of wavelength of scattered photon
13 printf ("\n wavelength of scattered photon is
     Angstrom. ", lambda1 * 1e10)
```

Scilab code Exa 3.21 Calculation of atomic number of unknown substance by Mosleys

```
1 clc
2 //Given that
3 lambda1 = 1.321 // wavelength of L- alpha line for
     platinum
4 lambda2 = 4.174 // wavelength of l - alpha line of
     unknown substance
5 z1= 78// atomic number of platinum
6 c = 3e8 // speed of light in m/s
7 b = 7.4 // constant for L - alpha line
8\ // \, \mathrm{Sample} Problem 21 page No. 146
9 printf("\n\n" Problem 21 # \n")
10 printf("\n Standard formula Used \n sqrt(nu1) = a*(Z-
     b)")
11 z^2 = b + (z^1 - b) * sqrt(lambda1 / lambda2) //
     calculation of the unknown substance has atomic
12 printf ("\n The unknown substance has atomic number
     \%d. ", z2)
```

 ${f Scilab\ code\ Exa\ 3.22}$ Calculation of zero point energy in box of length of 1 angstr

```
1 clc
2 //Given that
3 h = 6.6e-34 // plank's constant
4 m_e = 9.1e-31 // mass of electron in kg
5 L = 1e-10 // length of box of particle in m
```

```
6 //Sample Problem 22 page No. 146
7 printf("\n # Problem 22 # \n")
8 printf("\n Standard formula Used \n E= h^2 * (n_x^2+
      n_y^2+n_z^2) / (8*m*L^2)")
9 \quad sum = 0
10 \quad n_y = 1
11
        for n_x = 1:3
12
13
                  for n_z = 1:2
                      sum = n_x+n_y+n_z
14
15
                      if sum < 6 then
                            E = h^2 * (n_x^2+n_y^2+n_z^2)/
16
                                (1.6e-19*8*m_e*L^2) //
                                calculation of energy
                      \mbox{\tt printf}("\n\ \n\ E\%d\%d\%d\ is\ \ \%f\ eV. ",
17
                          n_x, n_y, n_z, E
18
                      end
19
20
21
22
             \quad \text{end} \quad
23
        end
```

Chapter 4

Frame of Reference

 ${
m Scilab\ code\ Exa\ 4.1\ Calculation\ of\ magnitude\ of\ force\ vector\ and\ it\ angle\ with\ axe}$

```
1 clc
2 //Given that
3 F = [2.5,4.5,-5] // F is a force vector act through
      origin
4 // sample Problem 1 Page No. 176
5 printf("\n\n" Problem 1 # \n")
6 \text{ F_magnitude} = \text{sqrt} (2.5^2 + 4.5^2 + (-5)^2)
7 theta_x = (180 / \%pi) * acos (2.5 / F_magnitude)
8 \text{ theta_y} = (180 / \%pi) * acos ( 4.5 / F_magnitude)
9 theta_z = (180 / \%pi) * acos ( -5 / F_magnitude)
10 printf (" \n Magnitude of force F is %f N",
      F_magnitude)
11 printf(" \n Angle made with X - axis is %f degree",
     theta_x)
12 printf(" \n Angle made with Y - axis is \%f degree",
     theta_y)
13 printf(" \n Angle made with Z - axis is %f degree",
     theta_z)
```

Scilab code Exa 4.2.1 To determine directional cosines of force vector

```
1 clc
2 //Given that
3 r = [2,2,2*sqrt(2)]
4
5 // sample Problem 2a Page No. 176
6 printf("\n \n\n # Problem 2a # \n")
7 r_magnitude = sqrt ( 2^2 + 2^2 + (2*sqrt(2))^2)
8 cos_x = ( 2 / r_magnitude)
9 cos_y = ( 2 / r_magnitude)
10 cos_z = ( 2.8284 / r_magnitude)
11 printf(" \n Directional cosine in X - axis is %f ", cos_x)
12 printf(" \n Directional cosine in Y - axis is %f ", cos_y)
13 printf(" \n Directional cosine in Z - axis is %f ", cos_z)
```

Scilab code Exa 4.2.2 To determine projection of a vector in xz and yz planes

```
1 clc
2 //Given that
3 r_xz = [2,2.8282]
4 // sample Problem 2b Page No. 176
5 printf("\n \n\n # Problem 2b # \n")
6 r_xz = sqrt (2^2 + (2.8282)^2)
7 r_yz = sqrt (2^2 + (2.8282)^2)
```

Scilab code Exa 4.3 Determination of work done by three forces

Scilab code Exa 4.4 Determination of vector to make airplane landing in time

```
1 clc
2 //Given that
```

```
3 v_w_x = 40 * \cos(45 * \%pi / 180) // x component of
      wind blow in miles/h
4 v_w_y = 40 * sin(45 * %pi /180) // y component of
      wind blow in miles/h
5 \text{ r_x} = 200 \text{ // distance} of destination point in x
      direction in miles
6 \text{ r_y} = 0 \text{ // distance} of destination point in y
      direction in miles
7 t = 40 // time taken by aeroplane to reach
      destination in minutes
8 // sample Problem 4 Page No. 177
9 printf("\n \# Problem 4 \# \n")
10 printf ("Standard formula used is V = V1 + V2 +
      \dots + V_n \setminus n ")
11 v_x = (r_x)/t *60 // x - component of velocity
      required to reach destination in time in miles/h
12 v_y = r_y / t *60 / x - component of velocity
      required to reach destination in time in miles/h
13 v_p_x = v_x - v_w_x // x component of aeroplane
      velocity in miles/h
v_p_y = v_y - v_w_y // y component of aeroplane
      velocity in miles/h
15 printf(" \n Vector of velocity of pilot with respect
       to moving air is %f i %fj miles/h \n where i and
       j stands for east and north respectively ",
     v_p_x,v_p_y)
```

 ${\bf Scilab}\ {\bf code}\ {\bf Exa}\ {\bf 4.5}\ {\bf Determination}\ {\bf of}\ {\bf difference}\ {\bf in}\ {\bf acceleration}\ {\bf due}\ {\bf to}\ {\bf gravity}\ {\bf at}$

```
1 clc  
2 //Given that  
3 R_e = 6.4e6 // radius of earth in m  
4 T = 8.64e4 // time period of one rotation of earth
```

```
5 theta_pole = 90 // angle between pole and rotational
       axis
6 theta_equator = 0 // angle between equator and
      rotational axis
7 g_pole = 9.8 // gravitational acceleration at pole
      in m/s^2
8 // sample Problem 5 Page No. 178
9 printf ("\n \n \ Problem 5 # \n")
10 printf ("Standard formula used is g1 = g - R_e * f^2 *
      \cos(\text{theta}))^2 \ 
11 f = 2 * \%pi / T // rotational frequency of earth
    g_{equator} = g_{pole} - R_{e} * f^2
13
    del_g = g_pole - g_equator
14 printf(" \n Difference in gravitational acceleration
       at pole and equator is \%e \text{ m/s}^2 ", del_g)
```

 ${f Scilab\ code\ Exa\ 4.6}$ Determination of angular velocity of earth if acceleration due

```
1 clc
2 //Given that
3 R_e = 6.4e6 // radius of earth in m
4 theta_pole = 90 // angle between pole and rotational
        axis
5 theta_equator = 0 // angle between equator and
        rotational axis
6 g_pole = 10 // gravitational acceleration at pole in
        m/s^2
7 g_equator = 0 // gravitational acceleration at
        equator in m/s^2
8 // sample Problem 6 Page No. 178
9 printf ("\n \n\n # Problem 6 # \n")
10 printf("Standard formula used is g1 = g - R_e*f^2*(
        cos(theta))^2 \n")
```

Scilab code Exa 4.7 Determination of angular velocity of earth if acceleration due

Scilab code Exa 4.8 Determination of magnitude and direction of Carioles force act

```
1 clc
2 //Given that
3 m = 1 // mass of particle in kg
4 theta = 30 // latitude position in degree
```

```
5 v = 0.5 // velocity of particle in km/s in north
      direction
6
7
8 // sample Problem 8 Page No. 180
9 printf("\n \# Problem 8 \# \n")
10 printf("Standard formula used is coriolis Force = 2*
     mass*angular velocity X velocity ")
11 f_x = -2*m*2*\%pi * v*1000*(-1)*sin(theta*\%pi/180)
     /86400 // coriolis force in east direction
12 	ext{ f_z = -2*m*2*\%pi * v*1000*cos(theta*\%pi/180)/86400}
     // coriolis force in verticle direction
13 F = sqrt(f_x^2+f_z^2)
14 alpha = -atan(f_z/f_x) *180 /\%pi
15 printf(" \n Magnitude and direction of coriolis
     force on particle are \n %e N and %d degree with
     east respectively", F, alpha)
```

Chapter 6

Relativistic Kinematics and Paradoxes in Relativity

 ${
m Scilab\ code\ Exa\ 6.1}$ Calculation in percentage contraction in length moving with sp

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 v = 0.8* c // velocity of rod
5 \ 11 = 1 \ // \ let
6 theta = 60 // anlge between length of rod and speed
     in degree
7 //Sample Problem 1 page No. 221
8 printf("\n \# Problem 1 \# \n")
9 l_x = 11 * cos(theta * %pi /180) * sqrt (1-(v /c)^2)
10 l_y = 11 * sin(theta * %pi /180)
11 12 = sqrt (1_x^2 + 1_y^2)
12 per_contrtaction = (11 - 12) / 11 *100
13 angle = atan (l_y/l_x)
14 printf ("Percentage contraction in rod is %f and
      apparant orientation is \%f", per_contrtaction, tan
      (angle) )
```

Scilab code Exa 6.2 Determination of relative velocity between two photons approach

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 \text{ u_x} = -3e8 // \text{ velocity of first photon in ground}
     frame \ in \ m/s
5 v = -3e8 // velocity of second photon in ground
     frame in m/s
6 // sample problem 2 page No. 222
7 printf("\n \n Problem 2 # \n")
8 printf("\n Standard formula used is u_x = (u_x + v)
     ) / (1 + v * u_x_ / c^2) ")
9 u_x = (u_x_+ v) / (1 + v * u_x_/ c^2) //
     calculation of Velocity of photon with respect to
10 printf ("\n Velocity of photon with respect to
     another is %d * c \setminus n Thus photons are approaching
      each other.",u_x / c)
```

 ${f Scilab\ code\ Exa\ 6.3}$ Determination of relative velocity between two spaceship approximation

```
5 v = -0.9 *c // velocity of second spaceship in
        ground frame in m/s
6 // sample problem 3 page No. 222
7 printf("\n \n\n # Problem 3 # \n")
8 printf("\n Standard formula used is u_x = (u_x_ + v
        ) / (1 + v * u_x_ / c^2) ")
9 u_x = (u_x_ + v) / (1 + v * u_x_ / c^2) //
        calculation of Velocity of photon
10 printf ("\n Velocity of photon with respect to
        another is %f c.", u_x / c)
```

Scilab code Exa 6.4 Determination of mass consumed to obtain energy

```
1 clc
2 //Given that
3 E = 7.5e11 // Energy in kWh
4 c = 3e8 // speed of light in m/s
5 // sample problem 4 page No. 223
6 printf("\n \n\n # Problem 4 # \n")
7 printf("\n Standard formula used \n E = m*c^2")
8 m = (E *3.6e6) / c^2// calculation of Amount of mass consumed
9
10 printf ("\n Amount of mass consumed is %d kg.", m)
```

 ${
m Scilab\ code\ Exa\ 6.5}$ Determination of energy that can be produced by consumption of

```
1 clc
2 //Given that
```

```
3 m = 4 // mass of substance consumed fully in kg
4 c = 3e8 // speed of light in m/s
5 // sample problem 5 page No. 223
6 printf("\n \n\n # Problem 5 # \n")
7 printf("\n Standard formula used \n E = m*c^2")
8 E = m * c^2// calculation of Amount of energy produced
9 printf ("\n Amount of energy produced is %e J.", E)
```

Scilab code Exa 6.6 Calculation of relativistic mass

Scilab code Exa 6.7 Determination of ratio of rest mass and relativistic mass of a

```
1 clc
2 //Given that
```

 ${\it Scilab\ code\ Exa\ 6.8.1}$ Determination of speed of space ship if observed length is h

 ${
m Scilab\ code\ Exa\ 6.8.2}$ Determination of time dilation if observed length is half of

```
1 clc
2 //Given that
```

Scilab code Exa 6.9 Calculation of mean life of meson moving with velocity

```
clc
//Given that
c = 3e8 // speed of light in m/s
v = 2.4e8 // speed of meson
t_0 = 2e-8 // lifetime of meson in second
// sample problem 9 page No. 224
printf("\n \n\n # Problem 9 # \n")
printf("\n Standard formula used ")
t = t_0 / sqrt(1 - (v / c )^2 ) // calculation of
Lifetime of meson
printf ("\n Lifetime of meson is %e sec.", t )
```

Scilab code Exa 6.10 Determination of velocity of 1 amu mass if it has kinetic ene

 ${f Scilab\ code\ Exa\ 6.11}$ Determination of velocity of a mass if it has total energy tw

Scilab code Exa 6.12.1 Determination of relativistic mass velocity observed by one

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 \text{ u_x} = -2e8 // \text{ velocity of first photon in ground}
     frame in m/s
5 v = -2e8 // velocity of second photon in ground
     frame in m/s
6 \text{ m}_{0} = 3e-25
7 // sample problem 12 page No. 226
8 printf("\n \n \ Problem 12a # \n")
9 printf("\n Standard formula used \n u_x = (u_x + v)
      / (1 + v * u_x_ / c^2)")
10 u_x = (u_x_+ v) / (1 + v * u_x_/ c^2) //
      calculation of Velocity of photon with respect to
      another
11 m = m_0 / sqrt(1 - (u_x / c)^2) / calculation of
      Relativistic mass of particle with respect to
     another
12 printf ("\n Velocity of photon with respect to
      another is \%e \text{ m/s.}", u_x)
13 printf ("\n Relativistic mass of particle with
     respect to another is %e kg.",m)
```

Scilab code Exa 6.12.2 Determination of relativistic mass observed by one mass for

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 ratio = 1.95e+03 // Ratio of relativistic mass and
    rest mass
5 // sample problem 14b page No. 227
6 printf("\n \n\n # Problem 14b # \n")
7 printf("\n Standard formula used \n m = m_o/ sqrt (
```

```
1- (v/c)^2)")
8 ratio_1 = 1 /(2* ratio^2) // calculation of ratio
    of velocity to velocity of light for
9 printf ("\n Ratio of velocity to velocity of light
    for particle is 1 - %e .", ratio_1)
```

Scilab code Exa 6.13 Determination of observed density of gold if it is moving wit

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 u = 0.9*c // velocity of first particle with respect
      to other in m/s
5 density1 = 19.3e-3 // density of gold in rest frame
6 // sample problem 13 page No. 226
7 printf("\n \n Problem 13 # \n")
8 printf("\n Standard formula used \n m = m_0 sqrt (
     1 - (v/c)^2 \n and \n l = l_o * sqrt ( 1- (v/c)^2)
9 mass_ratio = sqrt (1 - (u/c)^2) // calculation of
     ratio of relativistic mass
10 volume_ratio = 1 / sqrt (1 - (u/c)^2) //
     calculation of ratio of relativistic volume
11 density2 = density1 * (volume_ratio /mass_ratio ) //
     calculation of ratio of relativistic density
12 printf ("\n Relativistic density of rod in moving
     frame is %e.",density2)
```

Scilab code Exa 6.14.1 Determination of ratio of relativistic mass to rest mass

```
1 clc
2 //Given that
3 E = 1e9 // energy of electron in eV
4 c = 3e8 // speed of light in m/s
5 m_0 = 9.1e-31 // mass of electron in kg
6 // sample problem 14 page No. 227
7 printf("\n \n\n # Problem 14a # \n")
8 printf("\n Standard formula used \n E = m*c^2")
9 m = E / c^2 * 1.6e-19 // calculation of relativistic mass of particle
10 ratio = m / m_0// calculation of Ratio of relativistic mass and rest mass of particle
11 printf ("\n Ratio of relativistic mass and rest mass of particle is %e.", ratio )
```

Scilab code Exa 6.14.2 Determination of ratio of velocity of electron with respect

Scilab code Exa 6.14.3 Determination of ratio of their energy to rest mass energy

```
1 clc
2 //Given that
3 m = 9e-31 // mass in kg
4 E = 1e9 // Energy of accelerated electron in eV
5 c = 3e8 // speed of light in m/s
6 // sample problem 14c page No. 227
7 printf("\n \n\n # Problem 14c # \n")
8 printf("\n Standard formula used \n E = m*c^2")
9 E_0 = m * c^2// calculation of rest mass energy
10 ratio = E / E_0 *1.6e-19// calculation of Ratio of energy to rest mass energy
11 printf ("\n Ratio of energy to rest mass energy is %e.", ratio )
```

 ${f Scilab\ code\ Exa\ 6.15}$ Determination of proper length of rod if observed length is 1

```
9 l = l_ / sqrt (1 - (v / c)^2) // calculation of
Proper length of rod
10 printf ("\n Proper length of rod is %f m.",1)
```

Scilab code Exa 6.16 Determination of mean life of meson traveling with velocity

 ${f Scilab\ code\ Exa\ 6.17}$ Determination of velocity of electron having relativistic energy

```
1 clc
2 //Given that
3 E = 1 // energy of electron in MeV
4 c = 3e8 // speed of light in m/s
5 m_0 = 9e-31 // rest mass of electron
6 // sample problem 17 page No. 229
7 printf("\n \n\n # Problem 17 # \n")
```

Scilab code Exa 6.19 Calculation of distance travelled by meson

Scilab code Exa 6.20 Determination of speed if mass is increased by 1 percet

 ${\it Scilab\ code\ Exa\ 6.21}$ Determination of speed if mass is increased by 2000 times

Scilab code Exa 6.22 Determination of energy of each particle produced in pair pro

 ${\it Scilab\ code\ Exa\ 6.23}$ Determination of threshold wavelength for proton antiproton p

```
1 clc
2 //Given that
3 h = 6.63e-34 // plank's constant
4 c = 3e8 // speed of light in m/s
5 p_rest_mass = 938 // rest mass of proton in Mev/
```

 ${f Scilab\ code\ Exa\ 6.24}$ Determination of momentum of proton having kinetic energy 1Be

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 p_rest_mass = 0.938 // rest mass energy of proton in
      BeV
5 KE = 1 // kinetic energy of proton in BeV
6 // sample problem 24 page No. 232
7 printf("\n \n \ Problem 24 # \n")
8 printf("\n Standard formula used E^2 = p^2*c^2 + m_0
      ^2*c^4*")
9 E = KE + p_rest_mass// calculation of energy of
     particle
10 p = (sqrt (E^2 *1e6 - (p_rest_mass * 1e3)^2)) *(1.6e)
     -19)*(1e9) / c// calculation of Momentum of
     photon
11 printf("\n Momentum of photon is %e kg m/s.", p)
```

Scilab code Exa 6.26 Determination of speed of meson

 ${
m Scilab\ code\ Exa\ 6.27}$ Calculation of kinetic energy of a proton with velocity

```
1 clc
2 //Given that
3 c = 3e8 // speed of light in m/s
4 v = 0.8 *c // velocity of rod in m/s
5 m_0 = 1.673e-27 // rest mass of proton in kg
6 // sample problem 27 page No. 228
7 printf("\n \n\n # Problem 27 # \n")
8 printf("\n Standard formula used \n E_total = KE + E_mass")
```

Chapter 8

Our Solar System

Scilab code Exa 8.4.1 Comparison of speeds of two satellites

```
1 clc
2 //Given that
3 t1 = 1 // time period of satellite s1 in hours
4 t2 = 8 // time period of satellite s2 in hour
5 r1 = 1.2e4 // radius of orbit of satellite s1 in km
7 // sample problem 4a page No. 300
8 printf("\n\n" + Problem 4a # \n")
10 printf ("Standard formula r2/r1 = (t2/t1)^2(2/3)")
11 r2 = r1 * (t2/t1)^2(2/3) // calculation of radius of
      orbit of satellite s2 in km
12 v1 = 2 * %pi * r1 / t1 // calculation of speed of
     satellite s1 in km/h
                           // calculation of speed of
13 \text{ v2} = 2 * \%pi * r2 / t2
      satellite s2 in km/h
14 del_v = v2 - v1 // calculation of relative speed of
      satellites in km/h
15
```

```
16 printf (" \n Relative speed of satellite s2 wrt satellite s1 is %e km/h.", del_v)
```

Scilab code Exa 8.4.2 Comparison of angular speeds of two satellites

```
1 clc
2 //Given that
3 t1 = 1 // time period of satellite s1 in hour
4 t2 = 8 // time period of satellite s2 in hour
5 r1 = 1.2e4 // radius of orbit of satellite s1 in km
7 // sample problem 4b page No. 300
8 printf("\n\n\ Problem 4b # \n")
10 printf ("Standard formula r2/r1 = (t2/t1)^2(2/3)")
11 r2 = r1 * (t2/t1)^2(2/3) // calculation of radius of
     orbit of satellite s2 in km
12 v1 = 2 * %pi * r1 / t1 // calculation of speed of
     satellite s1 in km/h
13 v2 = 2 * \%pi * r2 / t2 // calculation of speed of
      satellite s2 in km/h
14 del_v = v2 - v1 // calculation of relative speed of
      satellites in km/h
15 del_r = r2 - r1 // calculation of closest distance
     between satellite s1 and s2
16 v_angular = del_v / del_r // calculation of angular
     speed in rad/h
17 printf (" \n Relative angular speed of satellite s2
     for satellite s1 is %e rad/h.", v_angular)
```

Scilab code Exa 8.5 Calculation of orbital velocity and period of revolution of sa

```
1 clc
2 //Given that
3 h = 2620 // distance of satellite from surface of
     Earth in km
4 R_e = 6400 // radius of Earth in km
5 M_e = 6e24 // mass of Earth in kg
6 G = 6.67e-11 // universal gravitational constant
8 // sample problem 5 page No. 300
9 printf("\n\n" Problem 5 # \n")
10
11 printf ("Standard formula used \n\t v_o = sqrt (G*M_e)
     r) \n ")
12 printf("\n \t T = 2 * pi * r / v_o \n")
13 r = R_e + h
14 \text{ v_o} = \text{sqrt}(G * M_e / (r * 1e3))
15 T = 2 * \%pi * r*1000 / (v_o*3600)
16 printf ("\n Orbital velocity of satellite is %f km/s
       \n period of revolution is \%f h.",v_o / 1000, T)
```

 ${f Scilab\ code\ Exa\ 8.6}$ Calculation of orbital velocity and period of revolution of sa

```
4 R_e = 6.38e6 // radius of Earth in km
5 \text{ M_e} = 6e24 // \text{ mass of Earth in kg}
6 g = 9.8 // gravitational acceleration in m/s2
8 // sample problem 6 page No. 301
9 printf("\n\n\ # Problem 6 # \n")
10
11 printf ("Standard formula used v_o = sqrt(G*M_e/r) \setminus n
12 printf("Standard formula used T = 2 * pi * r / v_o \
     n ")
13 r = R_e + h// calculation of effective distance
     between Earth and satellite
14
15 G = g * R_e^2 / M_e// calculation of gravitational
     constant
16 v_o = sqrt(G * M_e / r) / 1000 / calculation of
      orbital velocity of satellite
17 T = 2 * \%pi * r / (v_o * 1000) / 3.6e3 //
      calculation of period of revolution of satellite
18
19 printf ("\n Orbital velocity of satellite is %f km/s
       \n period of revolution is \%f h.", v_o, T)
```

Scilab code Exa 8.7 Estimation of mass of Earth

```
1 clc
2 //Given that
3 t = 27.3 // period of lunar orbit around Earth in
          days
4 r = 3.9e5 // distance of satellite from Earth in km
5 G = 6.67e-11 // universal gravitational constant
6 // sample problem 7 page No. 301
```

Scilab code Exa 8.8 Estimation of mass of sun

Scilab code Exa 8.9 Determination of height achieved by Rocket

```
1 clc
2 //Given that
3 R_e = 6.4e6 // radius of Earth in km
4 M_e = 6e24 // mass of Earth in kg
5 G = 6.67e-11 // universal gravitational constant
6 u = 6e3 // initial speed of rocket in m/s
8 // sample problem 9 page No. 302
9 printf("\n\n" Problem 9 # \n")
11 printf ("Standard formula used U_f - U_i = 1/2 * m *(
     u^2 - v^2 \setminus n ")
12 h = ((R_e * 1e3)^2 * u^2) / (2 * G * M_e - R_e * u)
     ^2) / 1000 // calculation of Height reached by
      rocket before returning to Earth
13
14 printf ("\n Height reached by rocket before
     returning is %e km.",h)
```

 ${f Scilab\ code\ Exa\ 8.10}$ Determination of velocity to be given to mass to achieve a parameter ${f Scilab\ code\ Exa\ 8.10}$

```
calculation of velocity required by mass to reach given height

12 printf ("\n Velocity required by mass is %e m/s.",v)
```

Scilab code Exa 8.11 Comparison of time period and speed of two planets

```
1 clc
2 //Given that
3 r1 = 1e12 // distance of first planet from Sun in m
4 r2 = 1e13 //distance of first planet from Sun in m
5 // sample problem 11 page No. 304
6 printf("\n\n" | Problem 11 # \n")
8 printf ("Standard formula used T^2 = k * r^3")
9 printf ("\n Standers formula used v = 2 * pi * r / T
10 r_ratio = r1 / r2 // r_ratio is ratio of distances
     from Sun
11 T_{\text{ratio}} = r_{\text{ratio}}(3/2) // \text{calculation of Ratio of}
     time period
12 v_ratio = r_ratio / T_ratio // calculation of ratio
     of speed
13
14 printf (" \n Ratio of time period is %f and ratio of
      speed is \%f .", T_ratio, v_ratio)
```

Scilab code Exa 8.12 Estimation of separation of Saturn from Sun

```
1 clc
```

```
2 //Given that
3 r1 = 1.5e8 // distance of Earth from Sun in km
4 t1 = 1 // let
5 // sample problem 12 page No. 305
6 printf ("\n\n\n # Problem 12 #\n")
7
8 printf ("\n Standard formula used T^2 = k* r^3")
9 t2 = 29.5 * t1// calculation of time period of Saturn
10 r2 = r1 * (t2 / t1) ^ (2/3) // calculation of distance of stern from Sun
11
12 printf ("\n Distance of Saturn from Sun is %e km ."
, r2)
```

 ${\it Scilab\ code\ Exa\ 8.13}$ Determination of speed of a satellite at perigee and apogee

```
1 clc
2 //Given that
3 r_peri = 360 // distance of perigee of satellite
    from Earth surface in km
4 r_apo = 2500 // distance of apogee of satellite from
    Earth surface in km
5 R_e = 6400 // radius of Earth in km
6 v_p = 30000 // speed of satellite at apogee position
    in km/h
7 // sample problem 13 page No. 305
8 printf ("\n\n\n # Problem 13 # \n")
9
10 printf ("\n Standard formula used v * r = k ")
11 r_p = r_peri + R_e // calculation of distance of
    perigee
12 r_a = r_apo + R_e// calculation of distance of
```

Scilab code Exa 8.14 Calculation of impulse magnitude and its direction required t

```
1 clc
2 //Given that
3 h = 600 // distance of satellite from surface of
      Earth in km
4 R_e = 6400 // radius of Earth in km
5 \text{ m\_s} = 100 \text{ // mass of satellite in kg}
6 \ g = 10 // gravitational acceleration in m/s2
7 v_y = 2500 // upward velocity of launched satellite
8 // sample problem 14 page No. 306
9 printf("\n\n\ Problem 14 # \n")
11 printf("\n Standard formula used 1/2 *(m_s * v ^2 /
      r) = g * R_E^2 * m / R_e^2 ")
12 r = R_e + h// calculation of effective height of
      satellite
13
14 \text{ v} = \frac{\text{sqrt}}{\text{g}} (\text{g} * (\text{R_e} * 1e3)^2 / (\text{r} * 1e3)) //
      calculation of orbital velocity of satellite
15
16 P_x = m_s * v// calculation of momentum in x
      direction
17 P_y = m_s * v_y// calculation of momentum in y
      direction
18 U = sqrt(P_x^2 + P_y^2) // calculation of magnitude
```

```
of impulse required

19

20 theta = (180 / %pi) * atan (P_y / P_x ) //
        calculation of direction of impulse required

21 printf ("\n Magnitude and direction of impulse
        required are respectively %ekgm/s and %f degree."
        ,U , theta)
```

Scilab code Exa 8.15.1 Calculation of loss of mass in the formation of 1 atom of h

```
1 clc
2
3 //Given that
4 b_e = 13.6 // Binding energy of electron to proton
    in eV
5 c= 3e8 // speed of light in m/s
6 // sample problem 15a page No. 306
7 printf("\n\n\n # Problem 15a # \n")
8 printf("\n Standard formula used E = m*c^2")
9 del_m = b_e * (1.6e-19) / c^2 * 1000
10 printf ("\n Loss in mass during formation of 1 atom
    of hydrogen is %e g.", del_m)
```

Scilab code Exa 8.15.2 Calculation of binding energy of deuteron

```
1 clc
2 //Given that
3 M_p = 1.6725e-24 // mass of proton in g
4 M_n = 1.6748e-24 // mass of neutron in g
```

```
5 M_d = 3.3433e-24 // mass of deuteron in g
6 c= 3e8 // speed of light in m/s
7 // sample problem 15b page No. 306
8 printf("\n\n\n # Problem 15b # \n")
9
10 printf("\n Standard formula used E = m*c^2")
11 del_m = M_p + M_n - M_d // calculation of Loss in mass during formation of 1 atom of hydrogen
12
13 b_e = (del_m / 1000) * c^2 / (1.6e-19 * 1e6) // calculation of Binding energy of deuteron
14
15 printf ("\n Binding energy of deuteron is %f MeV.", b_e)
```

Chapter 9

Stars and their Classification

Scilab code Exa 9.1 Calculation of change in brightness of a nova in 2 days

```
1 clc
2 //Given that
3 m_i = 15 // initial magnitude of supernova
4 m_f = 2 // final magnitude of supernova
5 // sample problem 1 page No. 332
6 printf("\n # Problem 1 # \n")
7
8 printf("Standard formula used \n\t M = m - 2.5 log(L/L_0)")
9 del_m = m_i - m_f // calculation of change in magnitude
10 brightness_ratio = 100^(del_m/5) // calculation of increment in brightness ratio.
11 printf ("\n In two days novas brightness is increased by %d times nearly", ceil(brightness_ratio / 10000)*10000)
```

Scilab code Exa 9.2.1 Calculation of change in magnitude if brightness get doubled

```
1 clc
2 //Given that
3 b_ratio = 2 // ratio of light output in a period
4 // sample problem 2a page No. 333
5 printf("\n # Problem 2a # \n")
6
7 printf("Standard formula used \n\t M = m - 2.5 log(L/L_0)")
8 del_m = 2.5 * log10(b_ratio) // calulation of change in magnitude
9 printf ("\n Change in magnitude is %f times", del_m
)
```

Scilab code Exa 9.2.2 Comparison of absolute brightness of Capella and sun

```
1 clc
2 // given that
3 m_capella = 0.05 // magnitude of brightness of
        capella at 14 parsecs
4 m_sun = 4.8 // absolute magnitude of brightness of
        sun
5 d = 14 // distance of capella in parsecs
6 D = 10 // distance of capella considerd for
        observation
7 // sample problem 2b page No. 333
8 printf("\n # Problem 2a # \n")
```